



Introductory and Refresher Course on Satellite and Lunar Laser Ranging



Lunar Laser Ranging

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History of LLR

How is LLR different from SLR

LLR network

LLR contribution to science

Challenges of LLR: next retroreflector's generation



LLR history

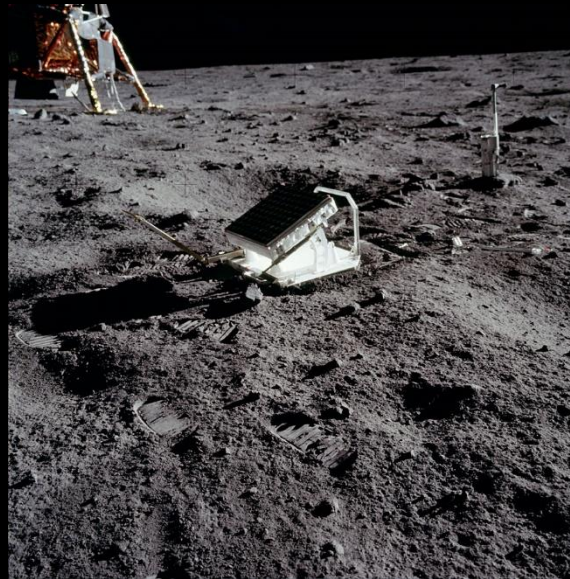


- Lunar laser ranging became possible after a retroreflector was placed on the Moon by the crew of Apollo 11.
- Five retroreflectors were placed on the Moon during the Apollo and Luna programs:
 - Apollo 11 in July 1969
 - Luna 17 (Lunokhod 1) in November 1970
 - Apollo 14 in February 1971
 - Apollo 15 in July 1971
 - Luna 21 (Lunokhod 2) in January 1973

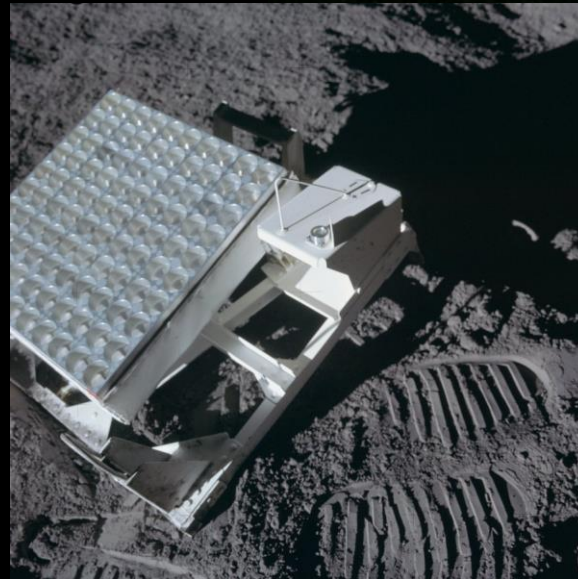
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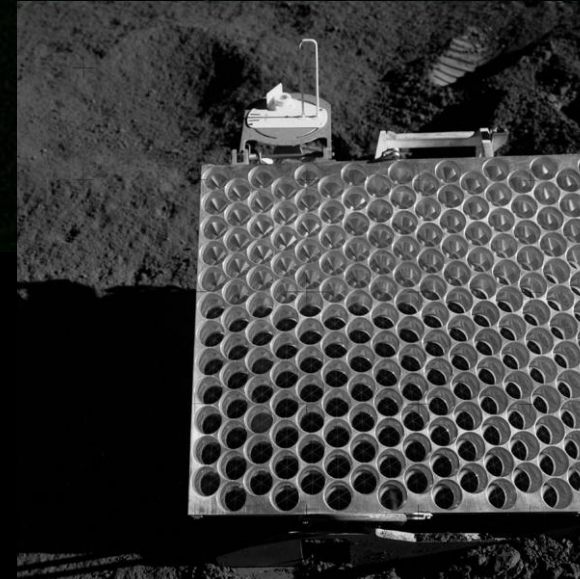
LLR history APOLLO retroreflectors



APOLLO 11
(07/1969)
Square 46 X 46 cm
100 corner cubes



APOLLO 14
(02/1971)
Square 46 X 46 cm
100 corner cubes



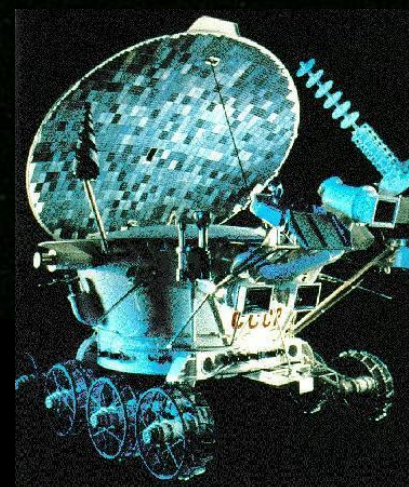
APOLLO 15
(07/1971)
Rectangle 104 X 61
cm
300 corner cubes



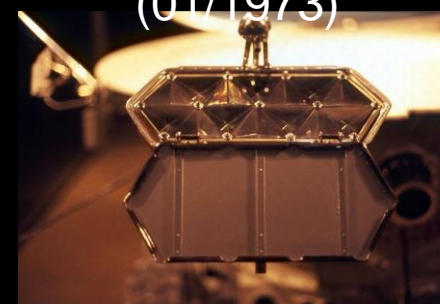
LLR history Lunokhod retroreflectors



Lunokhod 1
(11/1970)



Lunokhod 2
(01/1973)



Rectangle 44 X 19 cm
14 corner cubes



How is LLR different from SLR



- The link budget is function of power 4 of the distance:
 - SLR tracking: from 300km to 36 000km
 - LLR tracking: around 400 000km

- LLR requires more efficient equipment:
 - Larger telescope
 - More powerful laser
 - Better pointing and the tracking quality
 - Single photon detection

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Telescope



- The largest diameter at the emission: minimum divergence, small spot on the Moon. © Dan Long 2014
- The largest diameter at the reception: maximum number of photons.
- Good pointing: better than 1 arcsecond, but reference stars or craters can be used to correct the errors of the mount.
- Good tracking: better than 1 arcsecond for 10 minutes.



Laser



- The most powerful laser possible
 - The more narrow the pulsewidth, the less energy there is at the output.
- Due to the limited accuracy of the retroreflectors, and the weak link budget, short pulsewidth is not necessary (100ps).
- Ranging in infrared:
 - More energy
 - Less noise

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LLR network



- Four operational LLR stations:
 - APOLLO (USA), Grasse (France), Matera (Italy), and Wettzell (Germany).
- Stations in development:
 - In China, Russia, and South-Africa

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LLR contribution to science



- Five Retroreflectors Deployed
 - Apollo11, Luna17, Apollo14 and Apollo15 & Luna21 Missions
- Still Working
- Almost Daily Ranging Continues
- Analysis of Long Data History
- Evacuated Many Science Areas
 - Earth Science
 - Lunar Physics
 - Tests of General Relativity
 - Gravitation
 - Cosmology



RESULTS TO DATE



- Lunar Physics
 - Discover of Liquid Core – 15 years ago
 - Elastic Properties of the Crust

Earth Science Results

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Plate Tectonics

Question of Historical vs. Current Motion

LLRP has been Measuring the Current Motion

Earth Rotation

Evaluated the Changes in the Length of Day

Measurement of Polar Wander

Chandler Wobble to High Accuracy



EARTH SCIENCE RESULTS



- Plate Tectonics
 - Question of Historical vs. Current Motion
 - We Measured Current Motion
- Earth Rotation
 - Evaluated the Changes in the Length of Day
- Measurement of Polar Wander
 - Chandler Wobble to High Accuracy

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LLR GR RESULTS TO DATE



Equivalence principle parameter	η	$(6 \pm 7) \cdot 10^{-4}$
Metric parameter	$\gamma - 1$	$(4 \pm 5) \cdot 10^{-3}$
Metric parameter	$\beta - 1$: direct measurement	$(-2 \pm 4) \cdot 10^{-3}$
Time-varying gravitational constant	\dot{G}/G (year ⁻¹)	$(6 \pm 8) \cdot 10^{-13}$
Differential geodetic precession	$\Omega_{GP} - \Omega_{deSitter}$ (per century)	$(6 \pm 10) \cdot 10^{-3}$
Yukawa coupling constant	α (for $\lambda = 4 \cdot 10^5$ km)	$(3 \pm 2) \cdot 10^{-11}$
“Preferred-frame” parameter	α_1	$(-7 \pm 9) \cdot 10^{-5}$
“Preferred-frame” parameter	α_2	$(1.8 \pm 2.5) \cdot 10^{-5}$
Special relativistic parameters	$\zeta_1 - \zeta_0 - 1$	$(-5 \pm 12) \cdot 10^{-5}$
Influence of dark matter	$\delta_{galactic}$ (cm s ⁻²)	$4 \pm 4) \cdot 10^{-14}$

from Juergen Mueller and Franz Hofmann



GRAVITATIONAL & GR SCIENCE



- LLR Currently Provides our Best Tests of:

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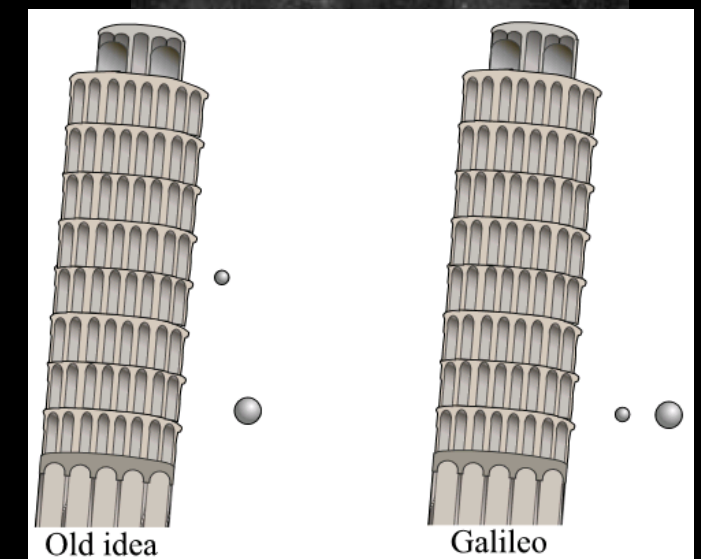
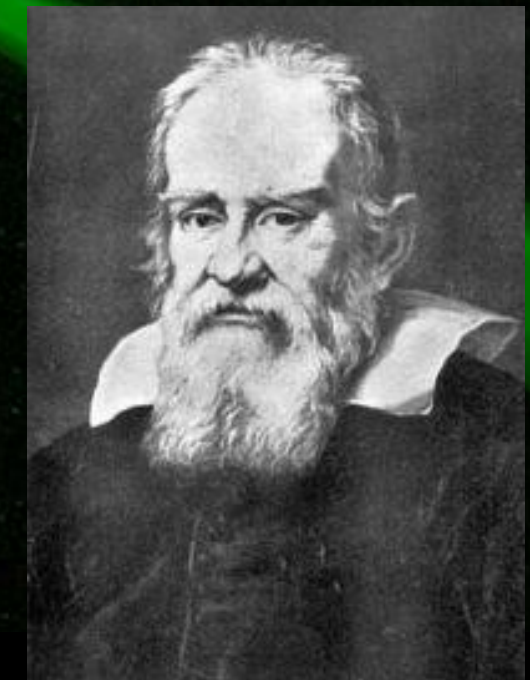
- The Strong Equivalence Principle (SEP)
- Time Rate-of-Change of G
- Inverse Square Law, Deviation of $1/r$
- Weak Equivalence Principle (WEP)
- Gravito-Magnetism



Weak Equivalence Principle



- Galileo's Apocryphal Experiment
 - Weak Equivalence Principle
 - Rate that the Earth and Moon Fall to the Sun
- Einstein is Correct © Dan Long 2014
 - In Absence of Air
 - All bodies Fall at Same Rate
 - Best Measurements to Date
- Even Gravity Energy is Hard to Push
 - Only Experiment to Measure This

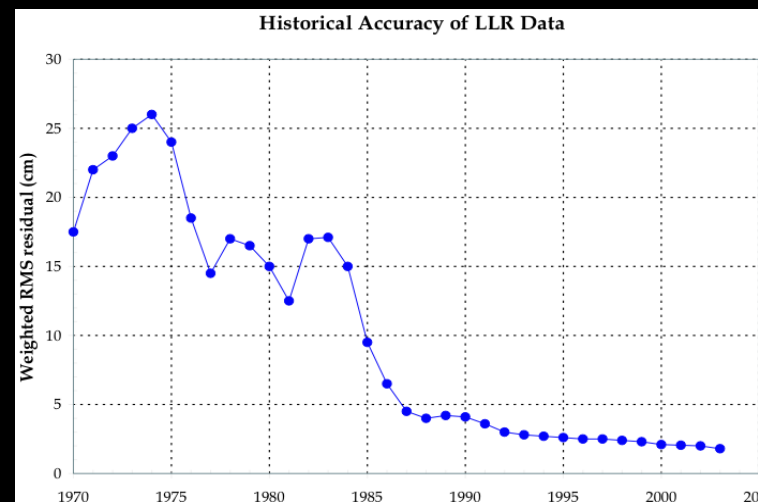




LIBRATION PROBLEM



- Why is there a Problem with the Apollo Arrays
 - Lunar Librations in Tilt Both Axis by 8/10
 - Apollo Arrays are Tilted by the Lunar Librations
 - Corner CCRs can have Different Ranges
 - As large as 100 mm for the Apollo 15 array
- Solution is One Large Retroreflector

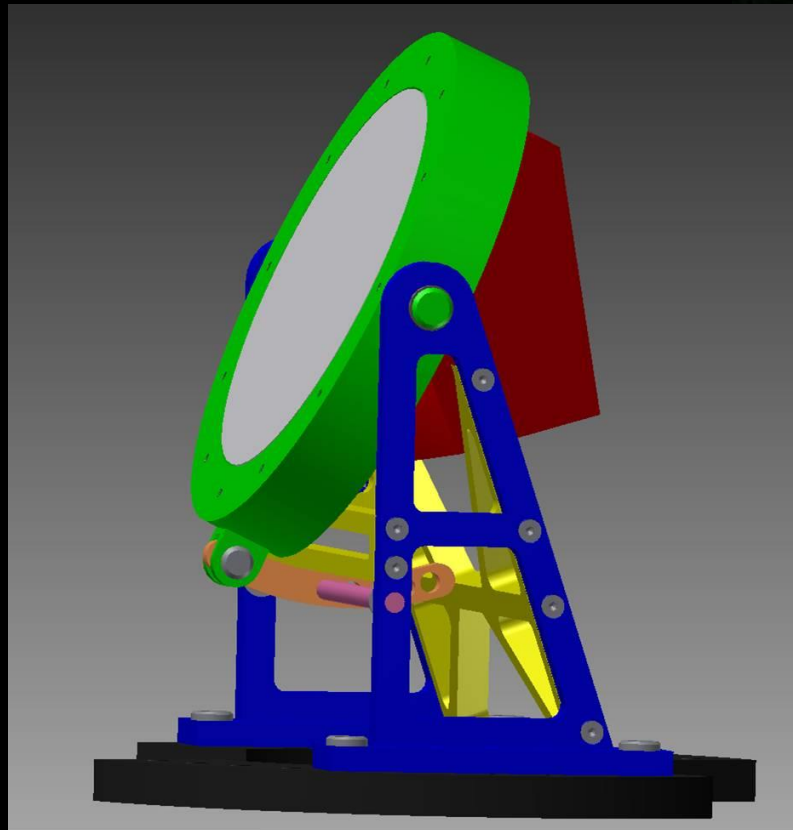




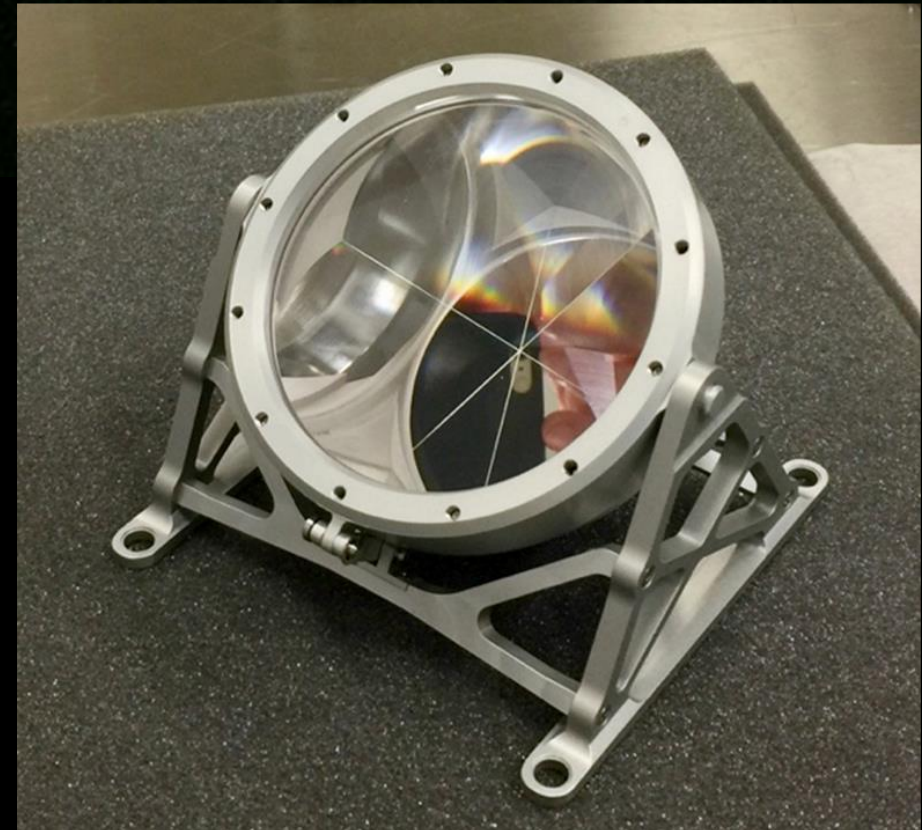
Current Next Generation RetroReflector



- Ranging Accuracy Improved by Up to a Factor of 100
- Limits to the Science Improvement
 - Ground Station Hardware and Procedures
 - Modeling of Horizontal Gradients in the Earth's Atmosphere



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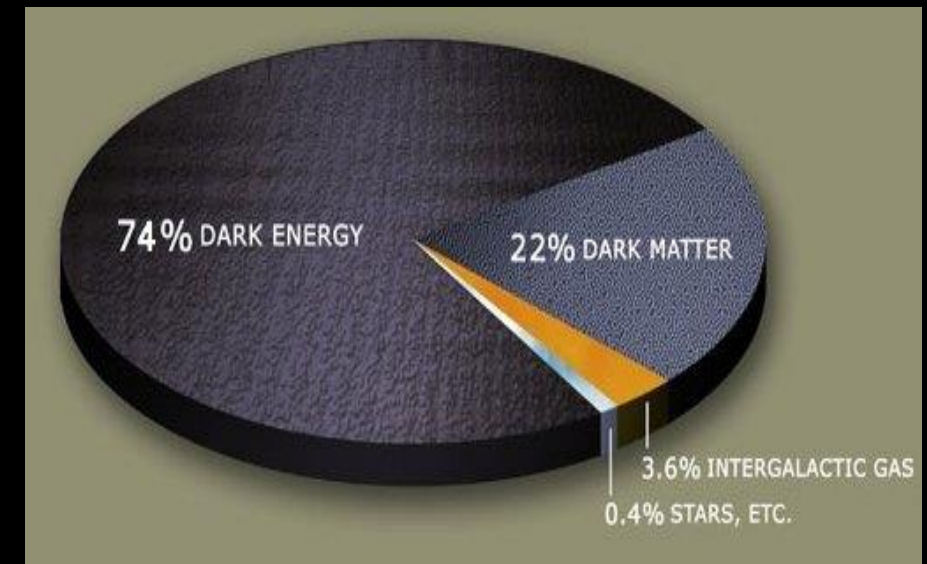


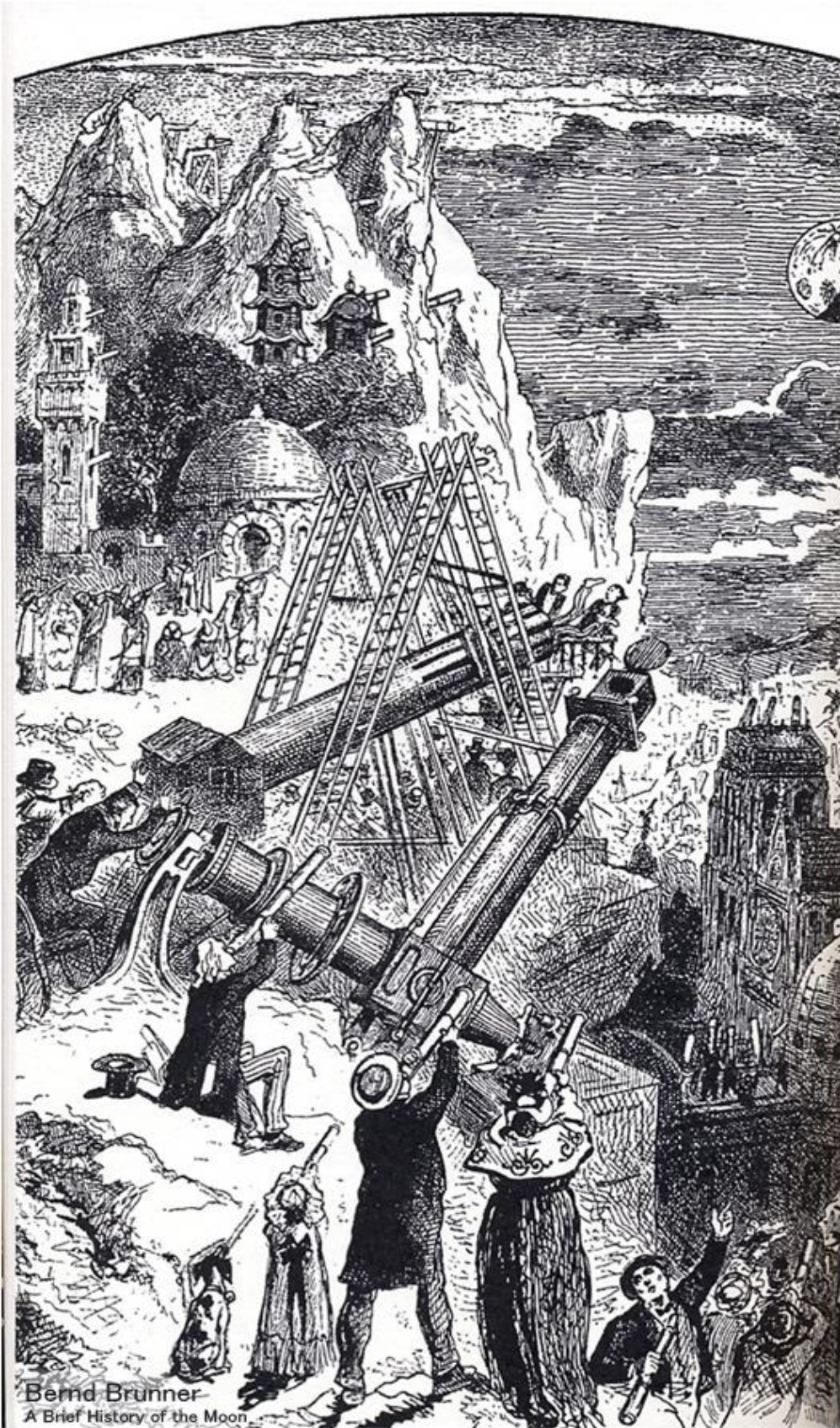


Conclusions



- Retroreflector Arrays Still Working after 50 years
- They Continue to Produce New Science
 - Lunar Science
 - Gravitational Physics and Tests of General Relativity
- NGLRs Will Improve Ranging Accuracy by up to 100
 - Limited only by Ground Stations and Atmospheric Modeling
- NASA Has Selected UMCP
 - To Deliver Three NGLRs
 - For Lunar Surface Deployment in 2021
- Why Push
 - 95% of Content of Universe is Unknown
 - GR and Quantum Mechanics in Conflict





Thank You!
any
Questions?
or
Comments?

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